

**AAS 01-327**

# **Trajectory Sensitivities for Sun-Mars Libration Point Missions**

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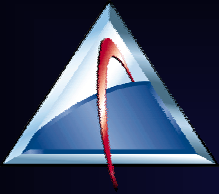
US Naval Postgraduate School

Monterey, CA

**AAS/AIAA Astrodynamics Specialists Conference**

**30 July – 2 August 2001**

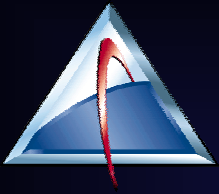
**Quebec City, Quebec, Canada**



# Overview



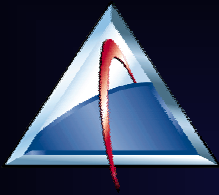
- Introduction
- Sun-Mars Libration Point Missions
- Mission Simulations and Analysis
- Targeting Methods using STK *Astrogator*
- Stationkeeping
- Conclusion



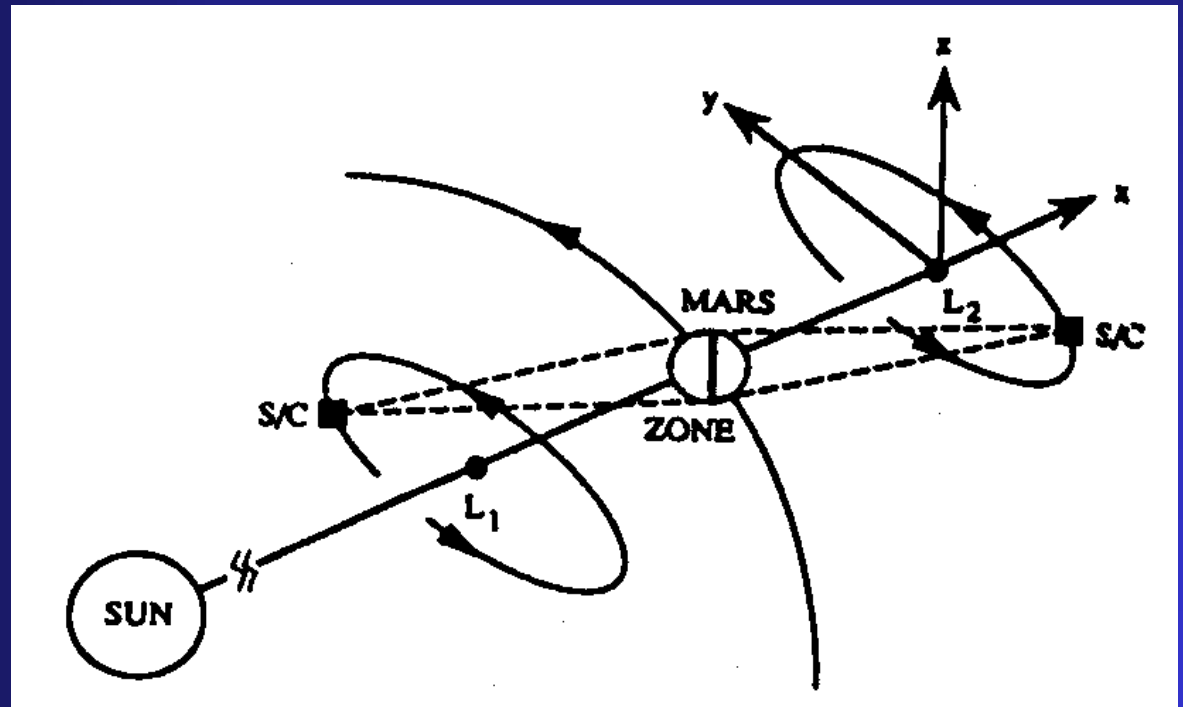
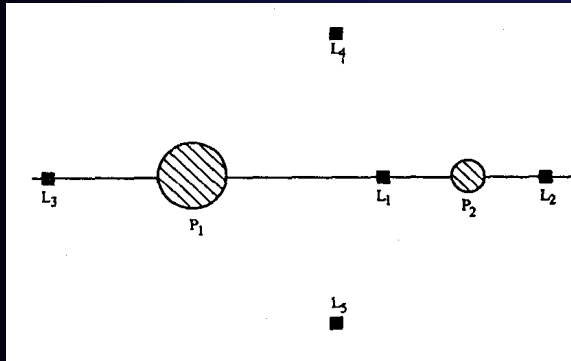
# Introduction

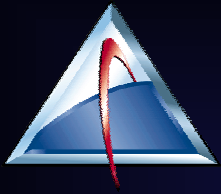


- *“NASA is seeking innovation to attack the diversity of Mars...to change the vantage point from which we explore...”* - CNN, 25 June 2001
- Sensitivities of 2016 Earth - Mars Transfer and LOI
- Desktop Computer Simulation
- Military Graduate Research / Industry Collaboration



# Sun-Mars Libration Point Missions

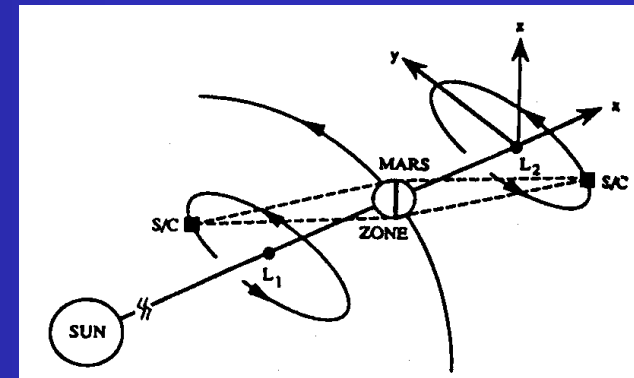


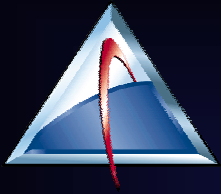


# Sun-Mars Libration Point Missions



- $L_1$ ,  $L_2$  Lagrange Points
  - 2 satellites – one in orbit about each point
  - Near-continuous coverage of Mars surface / orbit
  - Near-continuous link to Earth
- $L_1$ ,  $L_2$  Lissajous Orbits
  - Satellites orbit around  $L_1$ ,  $L_2$  points
  - Satellites opposed  $180^\circ$ , same direction orbit
  - Each satellite views “half” of Mars
    - ~99% of planet at all times

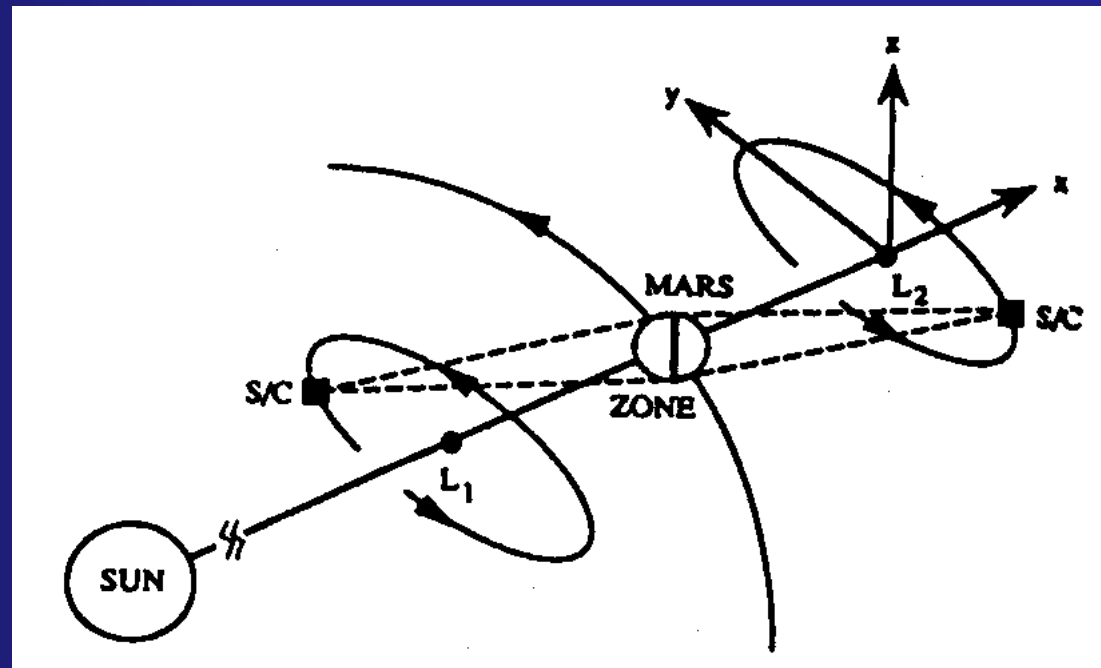


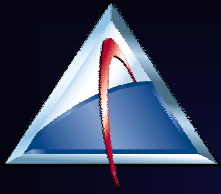


# Sun-Mars Libration Point Missions



- $L_1$ ,  $L_2$  Point Mission Considerations
  - Efficient Maintenance of  $180^\circ$  offset
    - insertion maneuvers, stationkeeping
  - Solar Exclusion Zone
    - Period  $> 0.9$  yrs

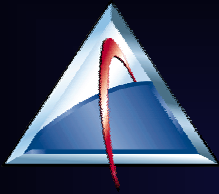




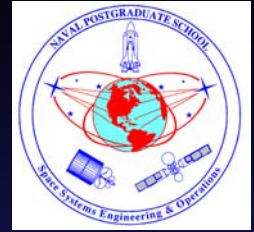
# Sun-Mars Libration Point Missions



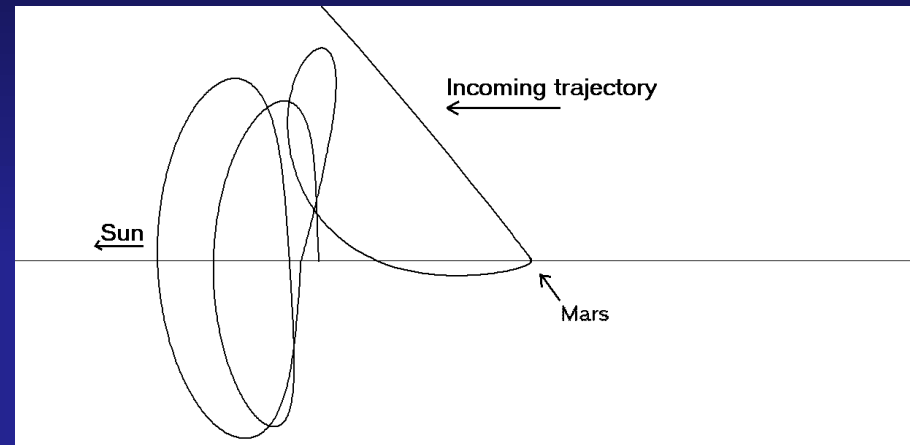
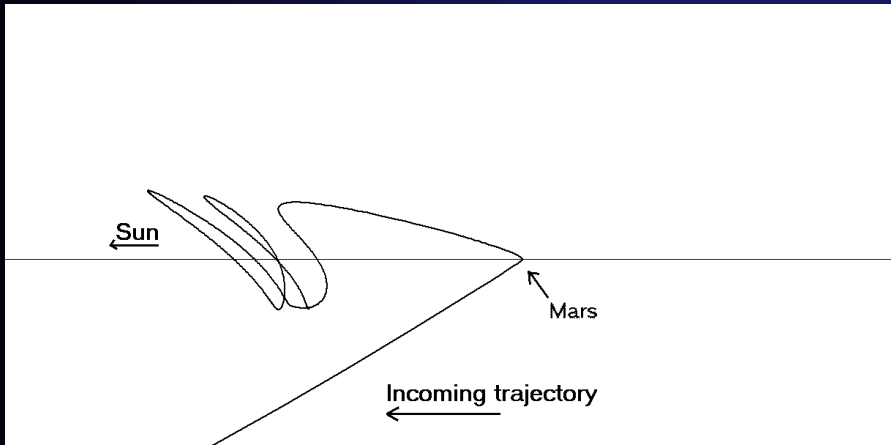
- Lissajous Orbit Constellation Advantages
  - 2 spacecraft required - minimum cost
  - L1 spacecraft can always see Sun, Earth
  - Long orbit period - simple tracking from Martian landers
  - Observation platforms
  - Small  $\Delta V$  maneuvers required
- Disadvantages
  - 1 million km distance - satellite to lander
  - Solar radiation interference
  - Loss of one satellite significant



# Mission Simulations and Analysis

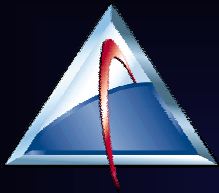


- 2003 Transfer with Braking Maneuver



Scenario	C3 Energy ( $\text{km}^2/\text{sec}^2$ )	Braking $\Delta v$ ( $\text{km}/\text{sec}$ )	Orbit Insertion $\Delta v$ ( $\text{km}/\text{sec}$ )	Total $\Delta v$ ( $\text{km}/\text{sec}$ )
Direct Injection	8.883	0	2.425	2.425
Braking Maneuver	9.056	0.856	0.104	0.960

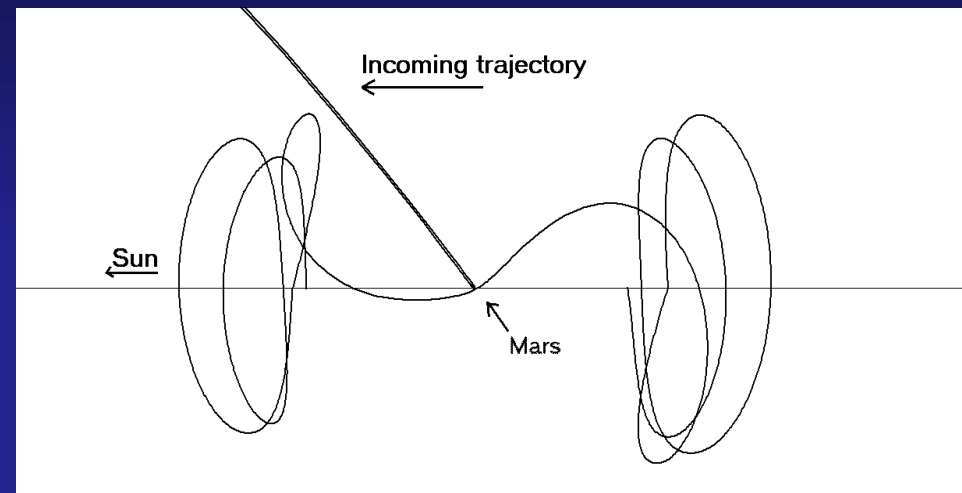
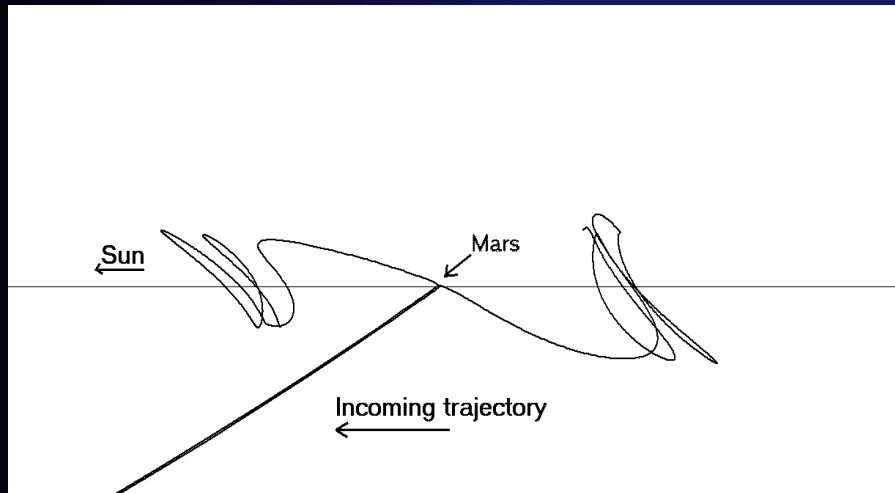




# Mission Simulations and Analysis

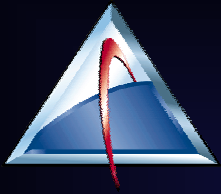


- 2016 Transfer with Braking Maneuver - Baseline



Orbit	C3 Energy ( $\text{km}^2/\text{sec}^2$ )	Mid-course $\Delta v$ (km/sec)	Braking $\Delta v$ (km/sec)	Orbit Insertion $\Delta v$ (km/sec)	Total $\Delta v$ (km/sec)
L1	10.377	0	1.710	0.047	1.757
L2	10.377	0.001	1.708	0.085	1.795

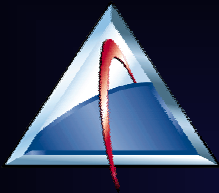
56 Day Difference Between LOI Epochs



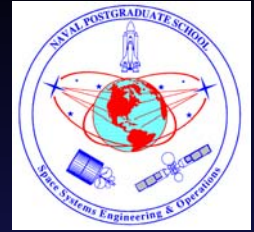
# Mission Simulations and Analysis



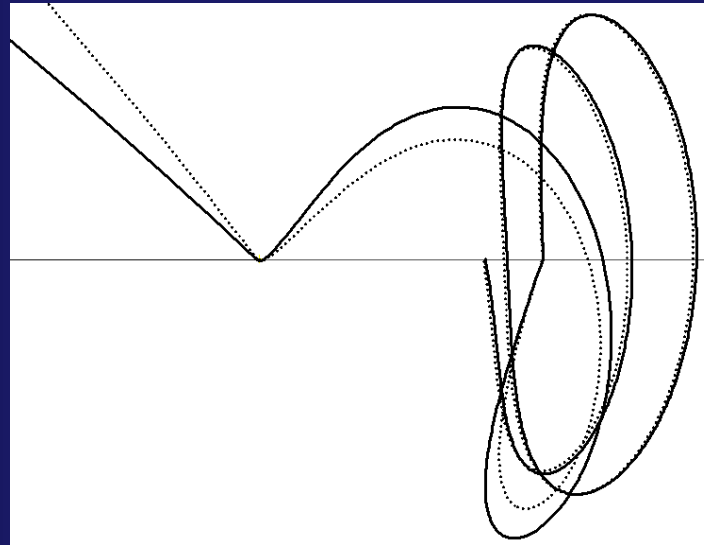
- Relative 180 deg Phasing Selection for S/C
- Achieve by:
  - Separate Launches
  - Relative phasing control via on-board propulsion
- Three methods to control phasing:
  - Midcourse Correction (MCC) Maneuver
  - TOF Adjustment from Mars Periapsis to LOI
  - Martian Phasing Loop



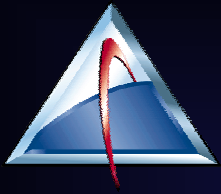
# Mission Simulations and Analysis



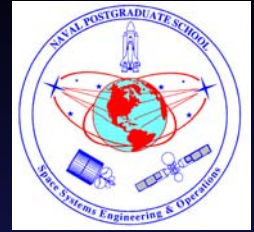
- Midcourse Correction (MCC) Maneuver
  - Change time of arrival at periapsis Mars, LOI
  - (Solid line)



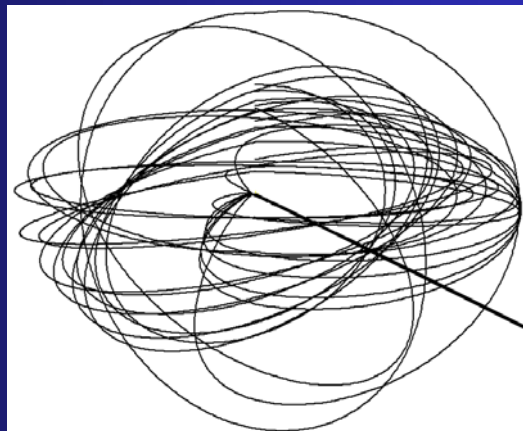
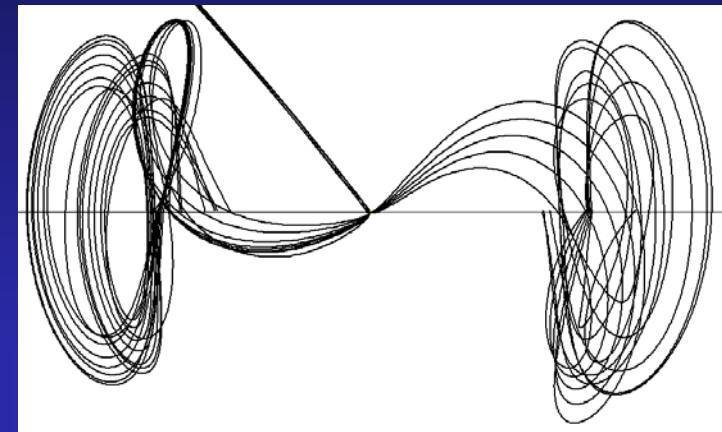
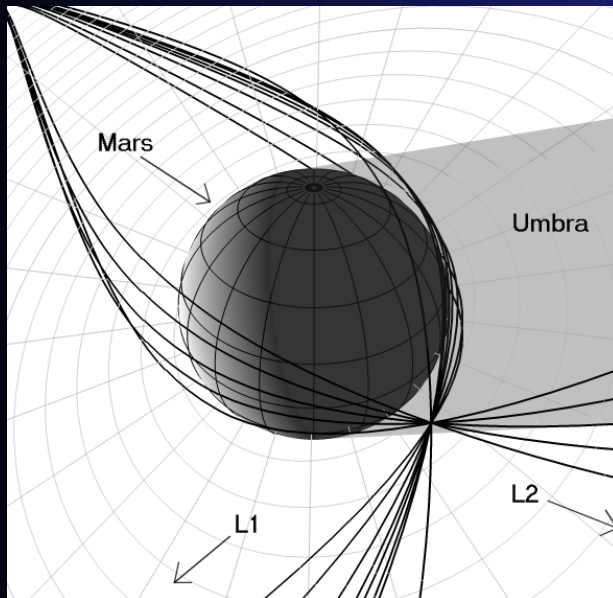
Vehicle/case	MCC Mag (km/s)	Retro $\Delta V$ (km/s)	Periapsis Date	LOI Diff. from L1 Orig. (Days)
L1 Original	0.00	1.71017	7 Sep 2016	0.00
L2 Original	0.00129	1.70843	7 Sep 2016	56.42
L2 -10 Days	0.281	2.00874	28 Aug 2016	58.62

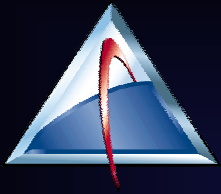


# Mission Simulations and Analysis



- TOF Adjustment from Mars Periapsis to LOI
  - B-plane correlates with Z-amplitude
  - Amplitude correlates with TOF

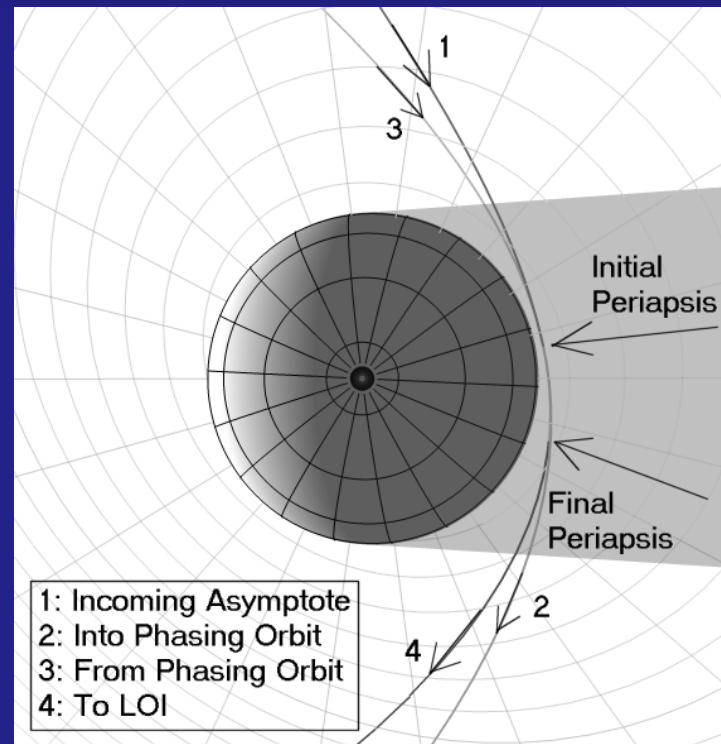
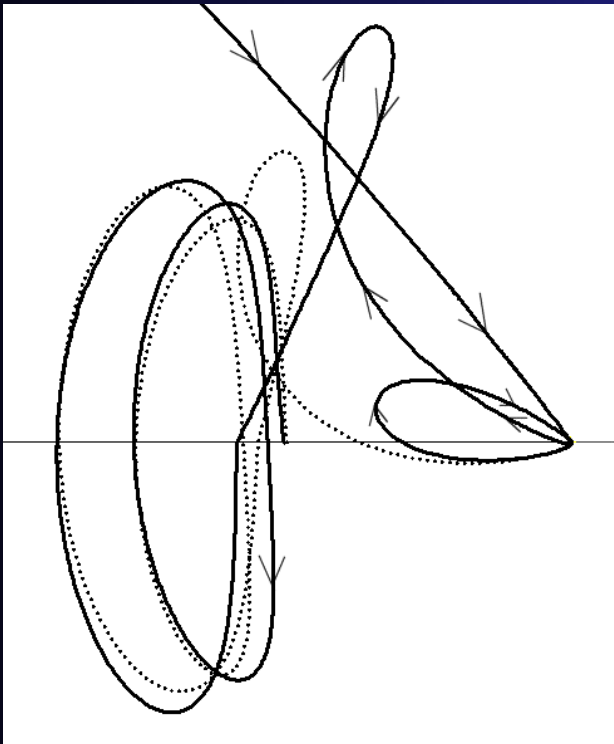


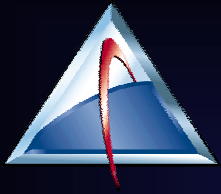


# Mission Simulations and Analysis



- Martian Phasing Loop
  - Phasing orbit period  $<$  LOI epoch difference due to periapsis rotation and transfer TOF





# Communication Coverage



- Use properly phased system to determine gaps

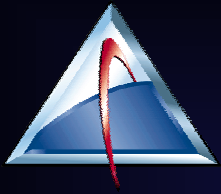
- Max revisit time

- duration of gap over interval

- Start at LOI

- propagate for 674 days
- determine visibility
- latitude points at one longitude

Latitude (deg)	Max Revisit Time (hrs)	Number of Gaps
90.0	0.000	1
80.0	3.639	61
70.0	5.117	335
60.0	3.499	584
50.0	1.707	902
40.0	0.696	890
30.0	0.494	846
20.0	0.441	785
10.0	0.463	733
0.0	0.486	688
-10.0	0.512	649
-20.0	0.543	609
-30.0	0.650	574
-40.0	0.845	524
-50.0	1.210	491
-60.0	5.994	436
-70.0	5.729	237
-80.0	7.425	130
-90.0	148.645	3

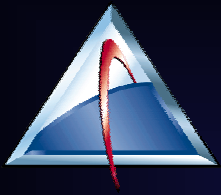


# Targeting Methods with STK Astrogator

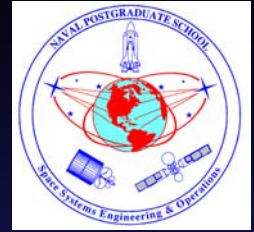


- Direct Transfer

Stage	Controls	Constraints	Dimension
I	C3 Targ.Vec. RA Targ.Vec. Dec	Epoch $X_{RLP}$ $Z_{RLP}$	3x3



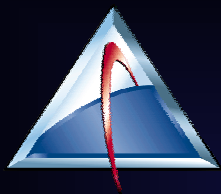
# Targeting Methods with STK Astrogator



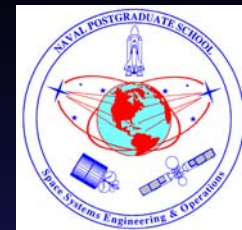
- Lissajous Orbit Insertion

Stage	Controls	Constraints	Dimension
I	$\Delta V_{LOlv}$ $\Delta V_{LOln}$ $\Delta V_{LOlc}$	Post LOI: $V_{xRLP}$ $V_{yRLP}$ $V_{zRLP}$	3x3
II	$\Delta V_{LOlv}$	1 <sup>st</sup> XZ Plane Cross: $V_{xRLP} = 0$	1x1
III	$\Delta V_{LOlv}$	2 <sup>nd</sup> XZ Plane Cross: $V_{xRLP} = 0$	1x1
IV	$\Delta V_{LOlv}$	3 <sup>rd</sup> XZ Plane Cross: $V_{xRLP} = 0$	1x1



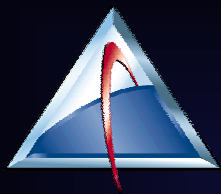


# Targeting Methods with STK Astrogator

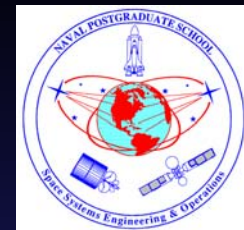


- Braking Maneuver

Stage	Controls	Constraints	Dimension
I	C3 Targ.Vec. RA Targ.Vec. Dec	Periapsis Epoch $B \cdot T$ $B \cdot R$	3x3
II	C3 Targ.Vec. RA Targ.Vec. Dec	Periapsis Epoch $B \cdot R$ $ R_p $	3x3
III	$\Delta V_{\text{retro}}$	1 <sup>st</sup> XZ Plane Cross: $X_{\text{RLP}}$	1x1
IV - VII	Same as LOI	Same as LOI	3x3, 1x1

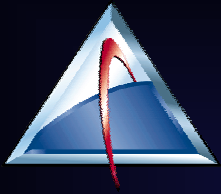


# Targeting Methods with STK Astrogator



- Z Amplitude Variations

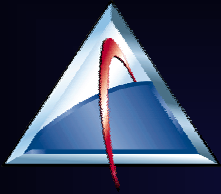
Stage	Controls	Constraints	Dimension
I	$\Delta V_{MCCx}$ $\Delta V_{MCCy}$ $\Delta V_{MCCz}$	Periapsis Epoch B·T B·R	3x3
II	$\Delta V_{MCCx}$ $\Delta V_{MCCy}$ $\Delta V_{MCCz}$	Periapsis Epoch B·R $ R_p $	3x3
Ila	Adjust B·R to get approximate $Z_{RLP}$ ; repeat stage II		
III	$\Delta V_{retro}$	1 <sup>st</sup> XZ Plane Cross: $X_{RLP}$	1x1
IV	$\Delta V_{MCCx}$ $\Delta V_{MCCy}$ $\Delta V_{MCCz}$ $\Delta V_{retro}$	Periapsis Epoch $ R_p $ 1 <sup>st</sup> XZ Plane Cross: $X_{RLP}$ $Z_{RLP}$	4x4
V - VIII	Same as LOI	Same as LOI	3x3, 1x1



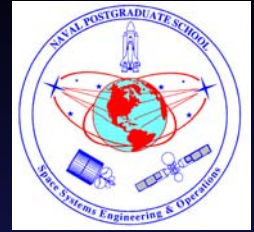
# Phasing Loop Targeting



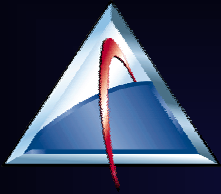
- Modified Target Procedure
  - After one phasing loop by L1 s/c, LOI matches L2 s/c
  - Two simultaneous differential corrector targeting schemes
- Inner Targeter
  - Transfer from phasing loop to LOI
- Outer Targeter
  - Retrograde maneuver at first Mars periapsis



# Stationkeeping



- Monte Carlo Analysis
  - Uncertainties modeled as uncorrelated errors
    - 100 m in position
    - 10 cm/s in velocity
    - 10% uncertainty in area of spacecraft
  - $\Delta V$  error of 10 cm/s
  - Vary these parameters; propagate L2 s/c for 90 days
  - SK maneuver to return trajectory to periodic; propagate for 1 year
  - Gather statistics for this “correction”
- 100 runs
  - Mars L2:  $DV = 0.044$  m/s (0.003 std dev)
  - Earth L2:  $DV = 0.45$  m/s (0.03 std dev)
  - Earth L2 (45 days):  $DV = 0.43$  m/s (0.03)



# Conclusion



- Three methods explored to control phasing
  - Use of phasing loop achieves 180 degree offset
- Communications coverage explored
  - Most latitudes experience ~30 minute gap
  - Poles experience few gaps, but longer (up to 6 days)
- Monte Carlo analysis for stationkeeping
  - Mars orbits require about order magnitude less than similar Earth orbit
  - Room for further study